

JPL

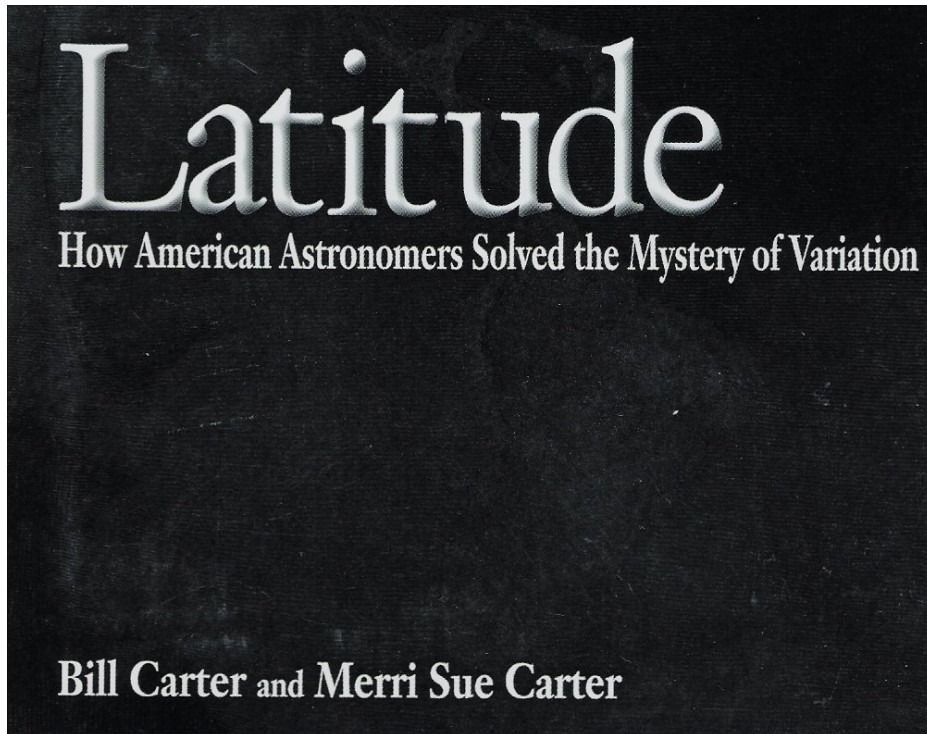
L. Romans

Lecture 4 – Spherical Trigonometry and related topics

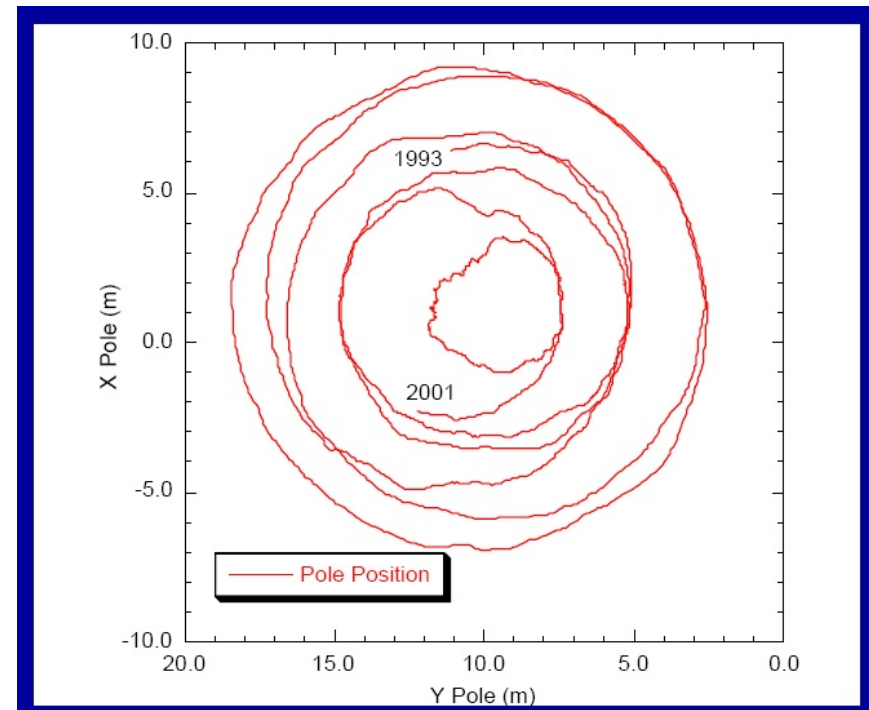


GISC-3325
24 January 2007

Another book recommendation



By Bill Carter and Merri Sue Carter,
Naval Institute Press, Annapolis,
Maryland 2002



Review

- Latitude and Longitude can uniquely and meaningfully describe where we are on the earth.
- We can also express positions *on a sphere* using 3-D Cartesian coordinates [X;Y;Z] using simple geometric relationships.

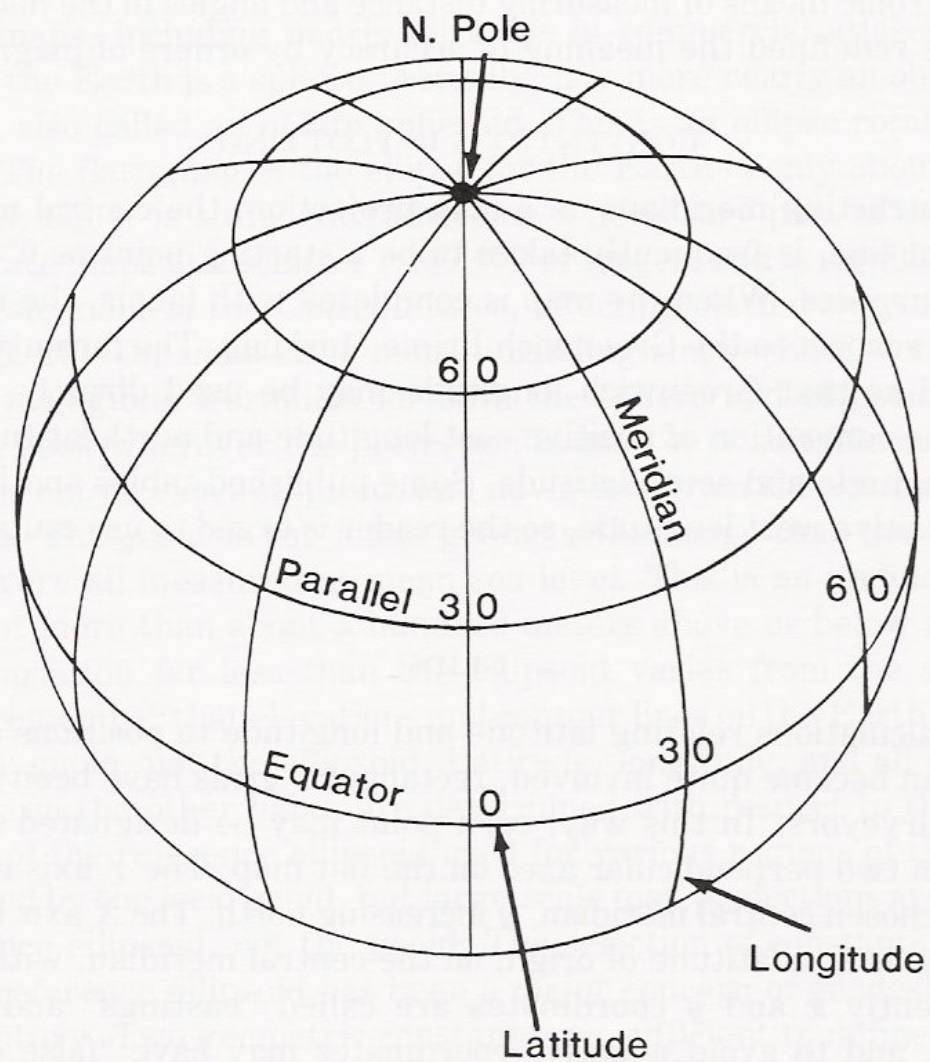
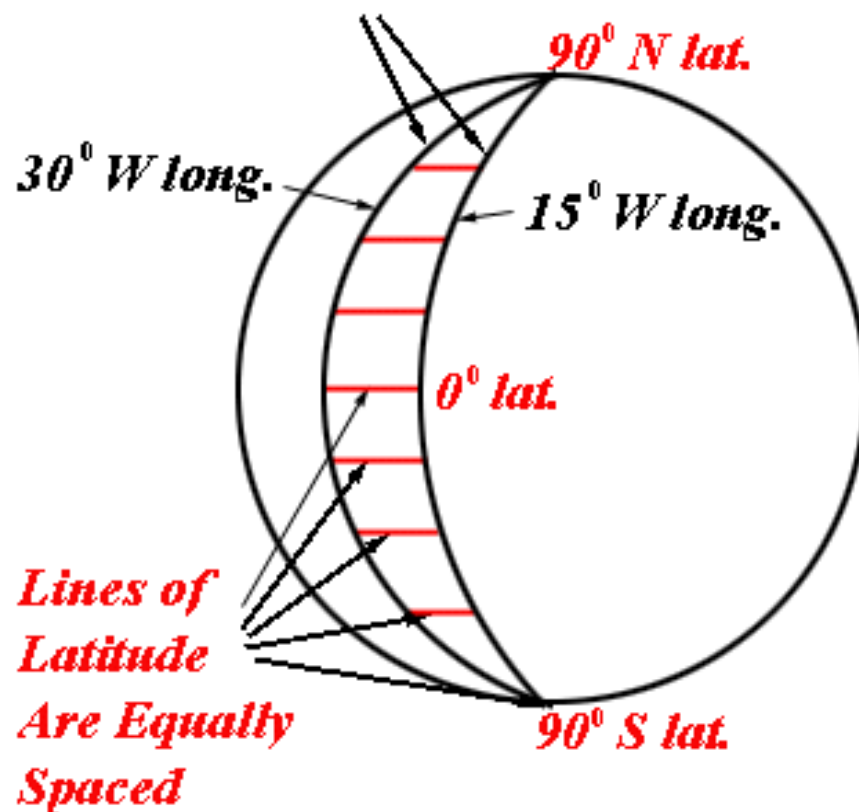


FIGURE 2.—Meridians and parallels on the sphere.

*Lines of Longitude
Converge*

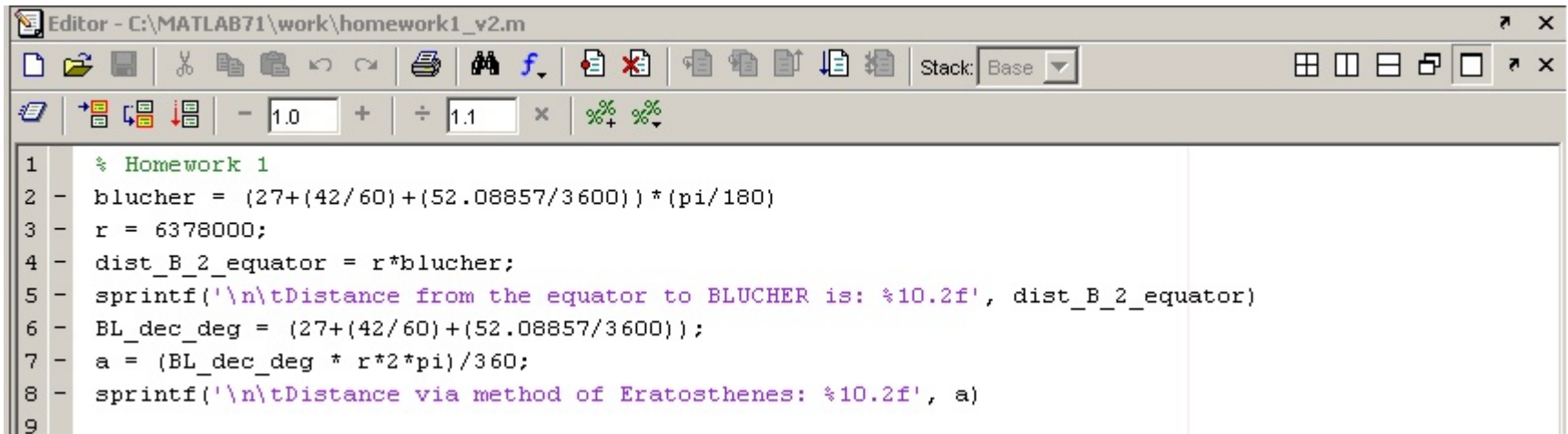


Textbook error

- See page 23.
- Author misplaces the decimal point where he converts DMS to Decimal
 - 31.315278 should be 3.1315278
- The answer to problem 2.2 is correct.

Homework Answer

- Problem: Use the position of station BLUCHER to determine the distance on a spherical earth (with radius 6,378,000 m) from the equator.
 - BLUCHER: 27-42-52.08857N

A screenshot of a MATLAB Editor window titled 'Editor - C:\MATLAB71\work\homework1_v2.m'. The window has a standard toolbar with icons for file operations, editing, and execution. Below the toolbar is a numeric keypad with fields for numbers 1.0 and 1.1, and buttons for mathematical operations. The main area of the window contains a MATLAB script with 9 lines of code. The script calculates the distance from the equator to station BLUCHER using two different methods. The first method uses the latitude of BLUCHER (27-42-52.08857N) converted to radians and multiplied by the Earth's radius (6,378,000 m). The second method uses the same latitude converted to degrees and multiplied by the Earth's radius and pi/180. Both methods yield the same result: 3,085,094 m.

```
1  % Homework 1
2  - blucher = (27+(42/60)+(52.08857/3600))*(pi/180)
3  - r = 6378000;
4  - dist_B_2_equator = r*blucher;
5  - sprintf('\n\tDistance from the equator to BLUCHER is: %10.2f', dist_B_2_equator)
6  - BL_dec_deg = (27+(42/60)+(52.08857/3600));
7  - a = (BL_dec_deg * r*2*pi)/360;
8  - sprintf('\n\tDistance via method of Eratosthenes: %10.2f', a)
9
```

Both methods yield: 3,085,094 m

What about using INVERSE?

Output from INVERSE

Ellipsoid : GRS80 / WGS84 (NAD83)
Equatorial axis, a = 6378137.0000
Polar axis, b = 6356752.3141
Inverse flattening, 1/f = 298.25722210088

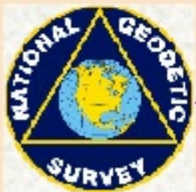
First Station : equator

LAT = 0 0 0.00000 North
LON = 97 19 44.31265 West

Second Station : BLUCHER

LAT = 27 42 52.08857 North
LON = 97 19 44.31265 West

Forward azimuth FAZ = 0 0 0.0000 From North
Back azimuth BAZ = 180 0 0.0000 From North
Ellipsoidal distance S = 3066800.0198 m



[NGS HOME PAGE](#)

3,066,800 meters

10,061,660 feet (US Survey)

1,905.6 miles (statute)

Our result using a radius of
6,378,000 meters is 3,085,094 m

A difference of 18,294 m!

Why??

UNITS ARE IMPORTANT

US v INTL feet

STATUTE v NAUTICAL MILES

Feet are Feet?

- Conversions from meters to feet (and inverse) are complicated by two units of feet.
 - U.S. Survey foot = 0.30480061... meters
 - $1200/3937$ meters (exactly)
 - International foot = 0.3048 meters (exactly)
 - 2.54 cm = 1 inch

DMS <-> Radian

- To convert degrees to radians
 - Convert DD MM SS.sssss to decimal
 - $\text{Deg} + \text{min}/60 + \text{sec}/3600$
 - Convert decimal degrees to radians
 - Multiply by $\pi/180$
- To convert radians to decimal
 - $\text{decDeg} = \text{Radian value} * 180/\pi$
 - $\text{Deg} = \text{floor}(\text{decDeg})$
 - $\text{Min} = \text{floor}((\text{decDeg} - \text{Deg}) * 60)$
 - $\text{Sec} = \text{decDeg} * 3600 - (\text{Deg} * 60) - (\text{Min} * 3600)$

The meter

- There were great difficulties in commerce due to varying length (and other) units.
- The French Academy of Science was charged with standardizing the measurement unit.
- Original proposal was to use the period of a pendulum.
- Instead, in 1790 the Academy recommended that a meter unit be based on one-millionth of the distance from the Equator to the North Pole.

How well did they do?

Output from INVERSE

Ellipsoid : Clarke 1866 (NAD27)
Equatorial axis, a = 6378206.4000
Polar axis, b = 6356583.8000
Inverse flattening, 1/f = 294.97869821380

First Station :

LAT = 0 0 0.00000 North
LON = 100 0 0.00000 West

Second Station :

LAT = 90 0 0.00000 North
LON = 100 0 0.00000 West

Forward azimuth FAZ = 0 0 0.0000 From North
Back azimuth BAZ = 180 0 0.0000 From North
Ellipsoidal distance S = 10001888.0430 m



[NGS HOME PAGE](#)

Output from INVERSE

Ellipsoid : GRS80 / WGS84 (NAD83)
Equatorial axis, a = 6378137.0000
Polar axis, b = 6356752.3141
Inverse flattening, 1/f = 298.25722210088

First Station :

LAT = 0 0 0.00000 North
LON = 100 0 0.00000 West

Second Station :

LAT = 90 0 0.00000 North
LON = 100 0 0.00000 West

Forward azimuth FAZ = 0 0 0.0000 From North
Back azimuth BAZ = 180 0 0.0000 From North
Ellipsoidal distance S = 10001965.7292 m

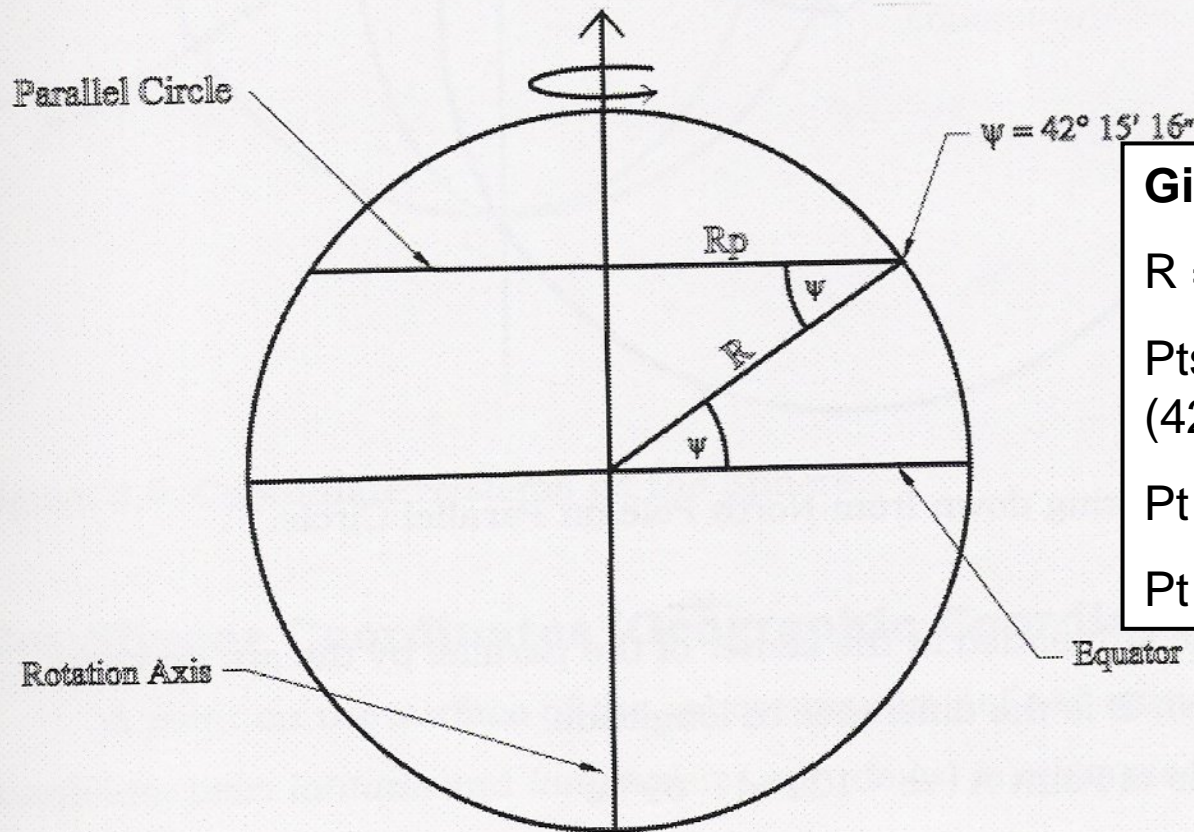


[NGS HOME PAGE](#)

Evolution of the meter

- The original measurement was in error due to unknown magnitude of Earth's flattening.
 - The unit was transferred to a platinum-iridium alloy bar kept in Paris (1874)
 - The unit was updated in (1889) to a bar composed of 90% platinum
- In 1960 a new definition was adopted that was based on krypton-86 radiation wavelength.
- Meter is the length of the path traveled by light *in a vacuum* during the time interval of:
 - $299\,792\,458\text{ s}^{-1}$ (299 792 458 meters per sec)

Distance on a sphere



Given

$$R = 6,371,000 \text{ m}$$

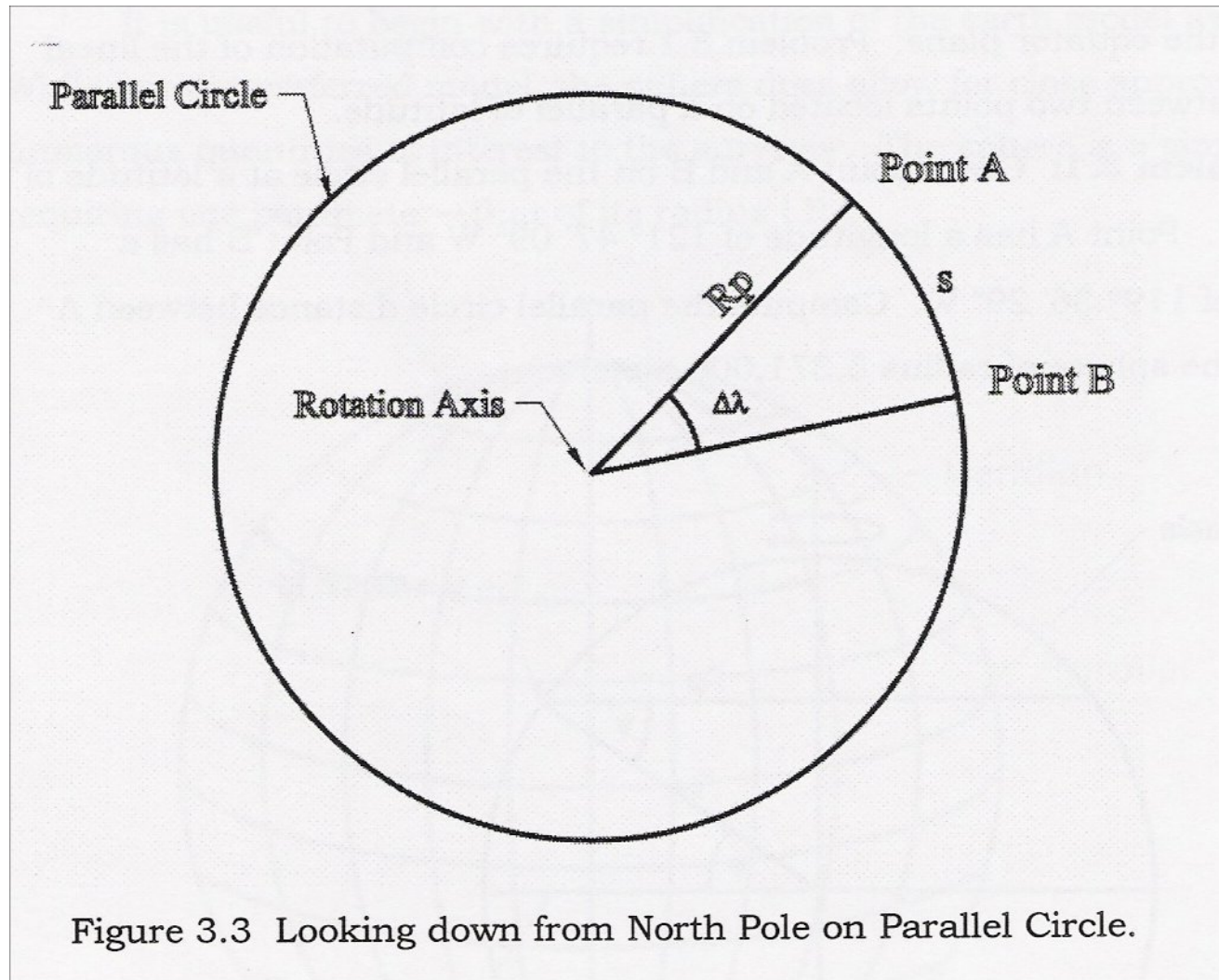
Pts A and B on same Latitude
(42d 15m 16s N)

Pt A Long: (121d 47m 09s W)

Pt B Long: (119d 36m 29s W)

Figure 3.2 Radius of Parallel Circle.

Compute radius of parallel circle by solving right triangle.



Subtract longitudes to get angle. $s = R_p * \text{angle (in radians)}$

For this problem $s = 179,237$ meters

Spherical Triangles

- Used in great circle navigation.
- Sides and angle are measured using arc measures
- Located on the surface of the sphere with sides formed by great circle arcs.
 - N.B. great circles are planes through the center of the Earth
- The shortest distance between points.
 - *Not exactly*

Spherical trigonometry

MAP PROJECTIONS—A WORKING MANUAL

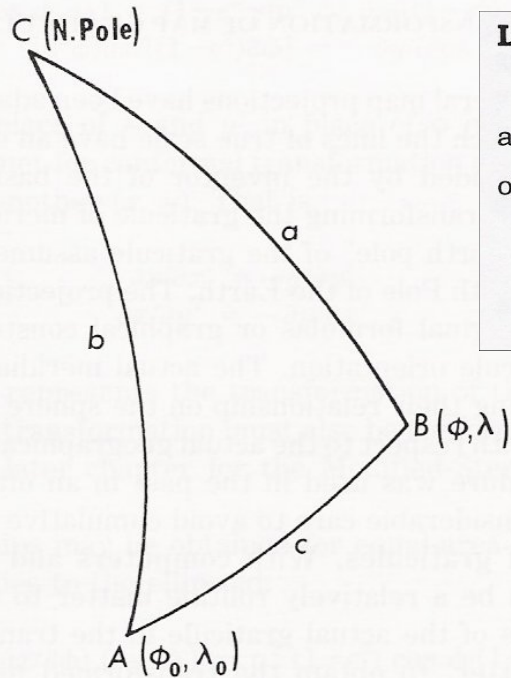


FIGURE 5.—Spherical triangle.

$$\cos c = \cos b \cos a + \sin b \sin a \cos C$$

Law of Cosines

There are two forms for the law of cosines - one for sides and one for angles. Note that the forms are both *cyclic*, i.e., a family of formulas may be obtained by cycling the side and angle variables.

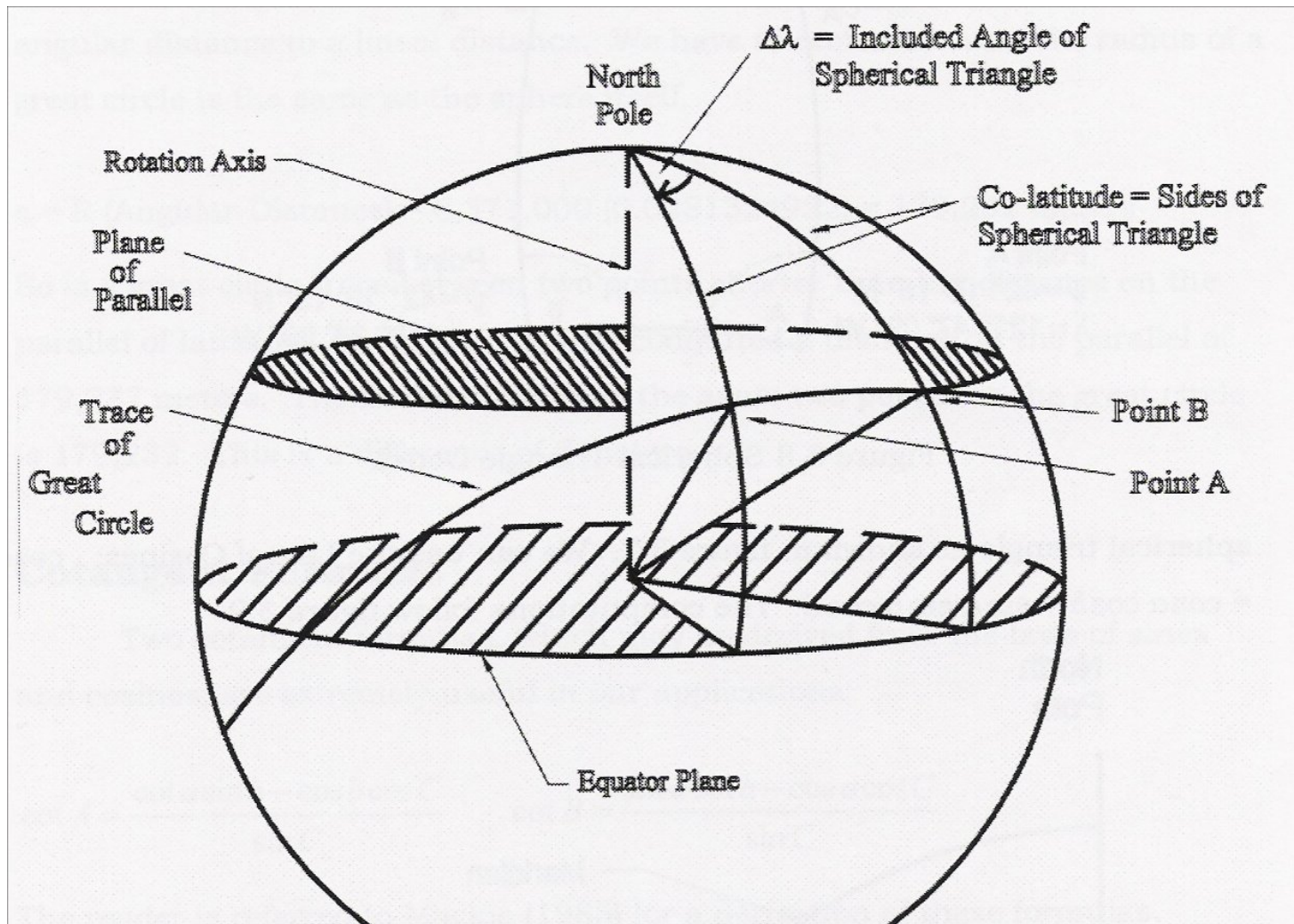
$$\cos a = \cos b \cos c + \sin b \sin c \cos A \quad (\text{law of cosines for sides}).$$

$$\cos A = -\cos B \cos C + \sin B \sin C \cos a \quad (\text{law of cosines for angles})$$

More damned definitions

- **Normal section** is a plane that contains the normal to the sphere at the occupied point and another point of interest.
- **Horizontal angle** is measured between two normal sections with respect to the instrument location.
- **Azimuth** is measured from the normal section containing the N pole clockwise to the normal section containing the other point
- **All normal sections on the sphere intersect the sphere along great circle arcs.**

Spherical Triangle



Problem in Text

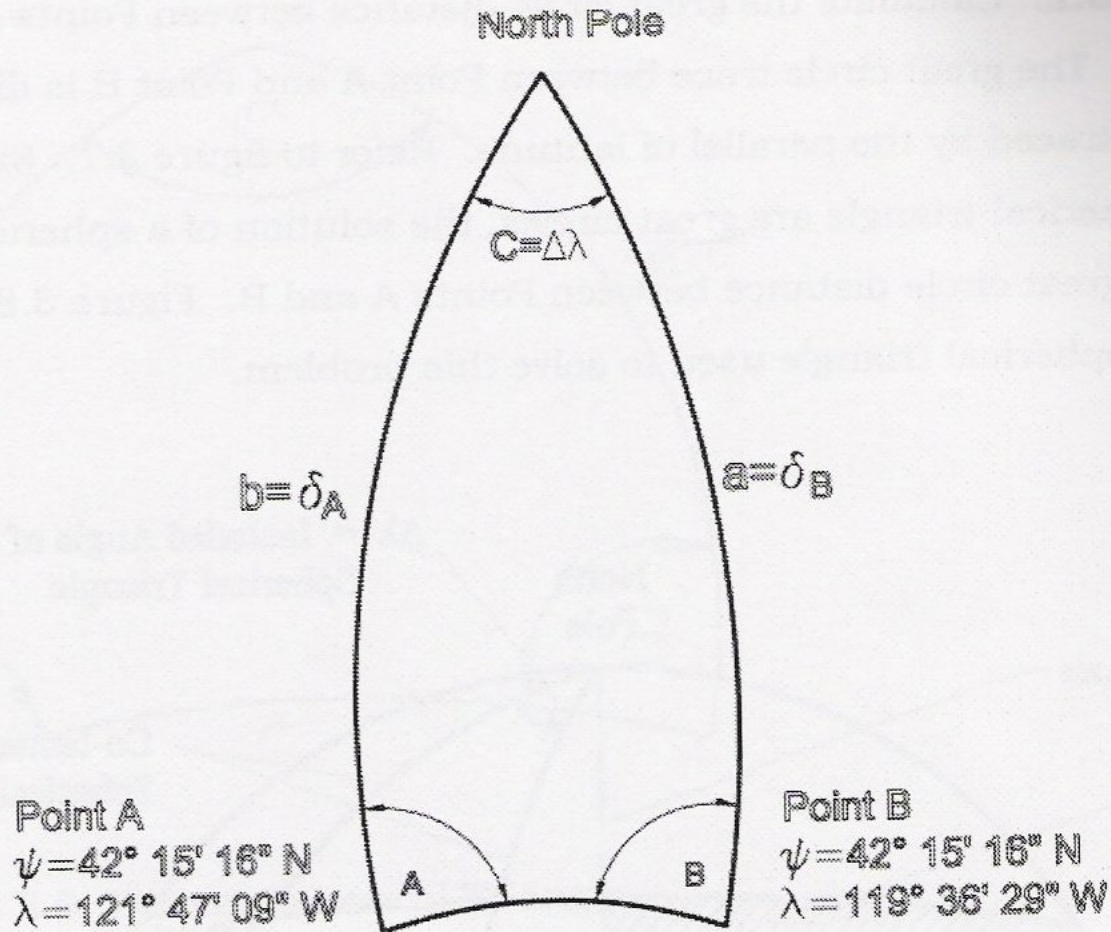


Figure 3.8 Spherical Triangle Detail.

Distance Calculations

- Determine the co-latitudes ($90^\circ - \text{latitude}$) at points A and B.
 - These are the lengths C to B (side a) and C to A (side b).
- Compute the difference in Longitudes
 - This is the angle at C
- As Law of Sines is ambiguous for angles in excess of 90° , we use Law of Cosines to solve for distance side c
 - $\cos(C) = \cos(A)\cos(B) + \sin(A)\sin(B)\cos(C)$.
- Distance = $r * C$ (N.B. radius of great circle same as the sphere itself).

We can also calculate Azimuth

- Use cotangent formulas and the results from the spherical triangle computation.
 - $\tan A = \sin C / ((\sin(b)/\tan(a)) - (\cos(b)\cos(C)))$
- Note that we must correctly account for the quadrant.
- Note as well that forward and reverse azimuths are not exactly 180 d different.

Other spherical Earth characteristics

- All meridians converge at poles.
- Azimuths of lines measured from one end to not equal values measured from the other end.
 - Effect is especially pronounced on long E-W lines.
- Can be approximated as a function of the E-W distance, mean latitude and spherical radius.

$$\theta'' = \frac{\rho \bar{d} \tan \bar{\psi}}{R}$$

Spherical excess

- The summation of all spherical angles exceed 180 degrees.
- It is proportional to the area of the spherical triangle.

$$\varepsilon = \frac{bc \sin A}{2R^2 \sin 1''}$$

```

#ap2006 1 1 0 0 0.00000000 96 ORBIT IGb00 HLM IGS
## 1356 0.00000000 900.00000000 53736 0.000000000000000
+ 29 1 2 3 4 5 6 7 8 9 10 11 13 14 15 16 17 18
+ 19 20 21 22 23 24 25 26 27 28 29 30 0 0 0 0 0
+ 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
+ 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
+ 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
++ 3 3 4 3 3 3 3 3 3 3 3 3 3 3 4 3 3
++ 4 3 3 3 3 3 4 3 3 4 3 3 0 0 0 0 0
++ 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
++ 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
++ 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
%c cc cc ccc ccc cccc cccc cccc cccc ccccc ccccc ccccc ccccc
%c cc cc ccc ccc cccc cccc cccc cccc ccccc ccccc ccccc ccccc
%f 0.00000000 0.0000000000 0.000000000000 0.0000000000000000
%f 0.00000000 0.0000000000 0.000000000000 0.0000000000000000
%i 0 0 0 0 0 0 0 0 0 0 0 0
%i 0 0 0 0 0 0 0 0 0 0 0 0
/* FINAL ORBIT COMBINATION FROM WEIGHTED AVERAGE OF:
/* cod emr esa gfz jpl mit ngs sio
/* REFERENCED TO IGS TIME (IGST) AND TO WEIGHTED MEAN POLE:
/* CLK ANT Z-OFFSET (M): II/IIA 1.023; IIR 0.000
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P 2 20133.793284 -12517.384408 11531.263733 -23.264373
P 3 -11366.447142 22473.599105 7788.691364 63.023929
P 4 26109.950844 -5714.160584 1216.379808 103.694692
P 5 -11013.232261 -21731.530250 -10967.998519 447.812380
P 6 -10108.117501 -15606.106217 19125.583219 170.870195
P 7 16494.168125 -8125.986115 -18697.734751 459.605468
P 8 26002.033551 50.071986 6220.016726 -51.940408
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P 10 5783.888427 -15080.250565 20946.780855 76.347737
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```

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BIT 2 OF LLI FLAGS DATA COLLECTED UNDER A/S CONDITION			COMMENT	
-Unknown-			MARKER NAME	
-Unknown-	-Unknown-		OBSERVER / AGENCY	
-Unknown-	ASHTech UZ-12	ZE21	REC # / TYPE / VERS	
-Unknown-	-Unknown-		ANT # / TYPE	
-726296.8700 -5598342.1900 2958518.6100			APPROX POSITION XYZ	
0.0000 0.0000 0.0000			ANTENNA: DELTA H/E/N	
1 1			WAVELENGTH FACT L1/2	
7 L1 L2 C1 P2 P1 D1 D2			# / TYPES OF OBSERV	
SNR is mapped to RINEX snr flag value [1-9]			COMMENT	
L1: 1 -> 1; 90 -> 5; 210 -> 9			COMMENT	
L2: 1 -> 1; 150 -> 5; 250 -> 9			COMMENT	
2005 9 14 13 37 15.00000000 GPS			TIME OF FIRST OBS	
			END OF HEADER	
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